



Incoherent Scatter Radar (ISR), Pulsating Aurora, & Ion Outflow

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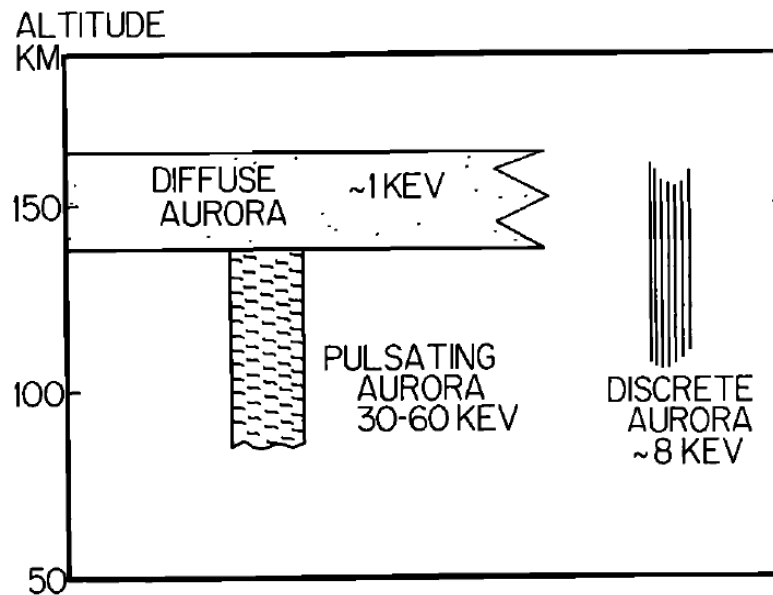
- Background
 - Pulsating Aurora
 - Ion Outflow
- ISR examples
 - Ion Upflow
 - Pulsating Aurora
- Motivation - e-POP observations
- Proposed experiment



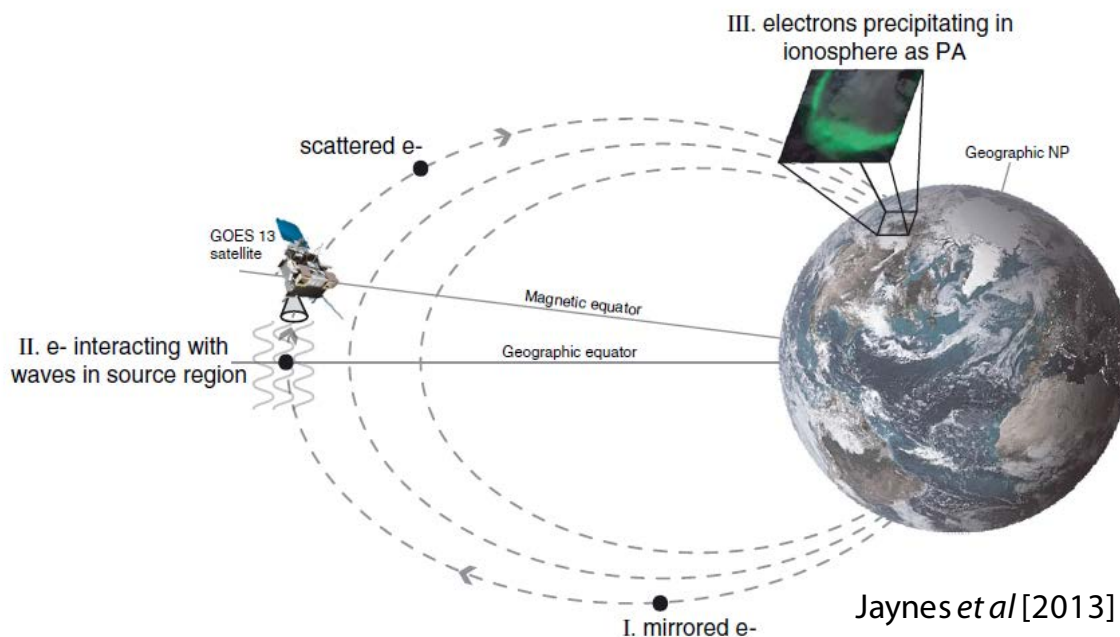
Pulsating Aurora



- Irregular shape
 - 10s - 100s km wide
 - Vertically thin
- Irregular period
 - 2-20 s
 - 8 s average
- Optically dim
 - ≈ 1 kR (427.8 nm)
- Consistent energy range

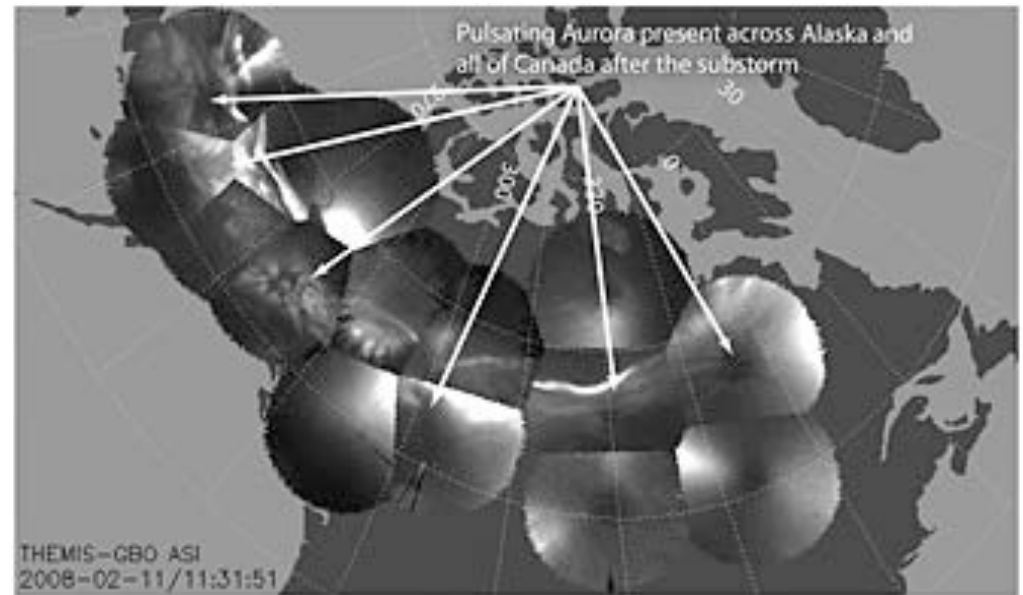


Brown *et al* [1976]

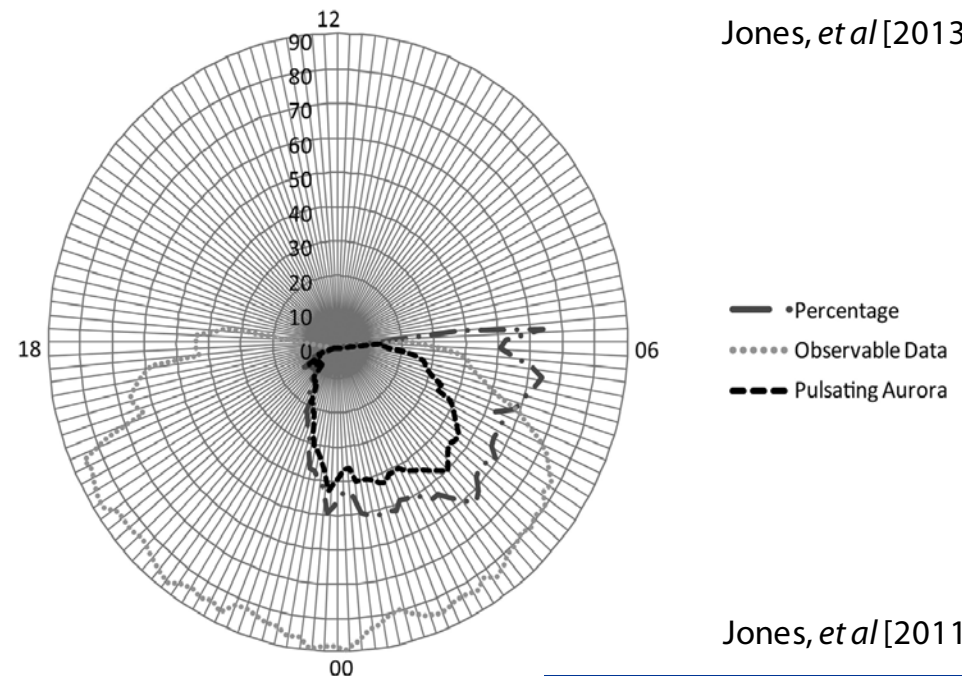


Jaynes *et al* [2013]

- Widespread temporally and geographically
- From Jones, *et al* [2011] statistical study:
 - Most probable duration 90-120 min
 - 31% clear optical data exhibit PA
 - 69% of PA occur post substorm
 - 54% probability to occur after magnetic midnight

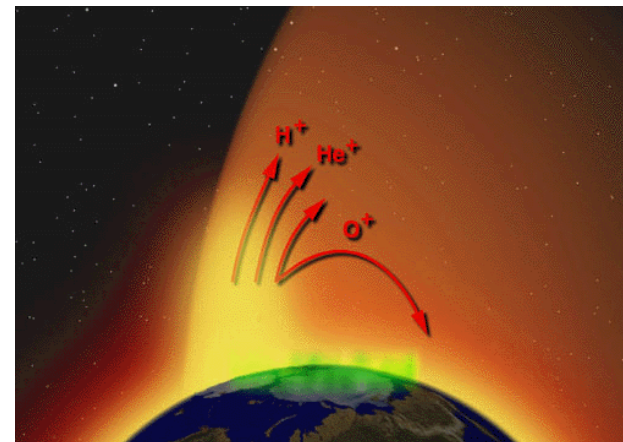
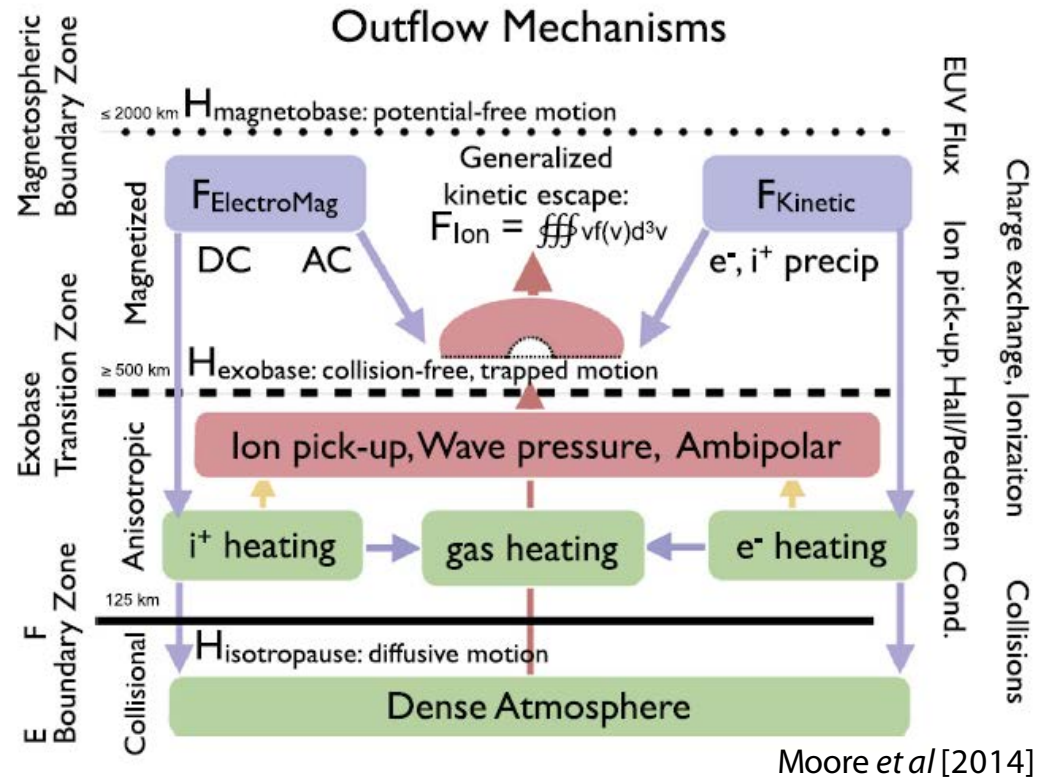


Jones, *et al* [2013]



Jones, *et al* [2011]

- Magnetosphere-Ionosphere coupled in many ways
 - Polar outflow
 - Particle precipitation
 - **Ion outflow**
- Addition of heavy ions (i.e. O^+) to magnetosphere from ionosphere

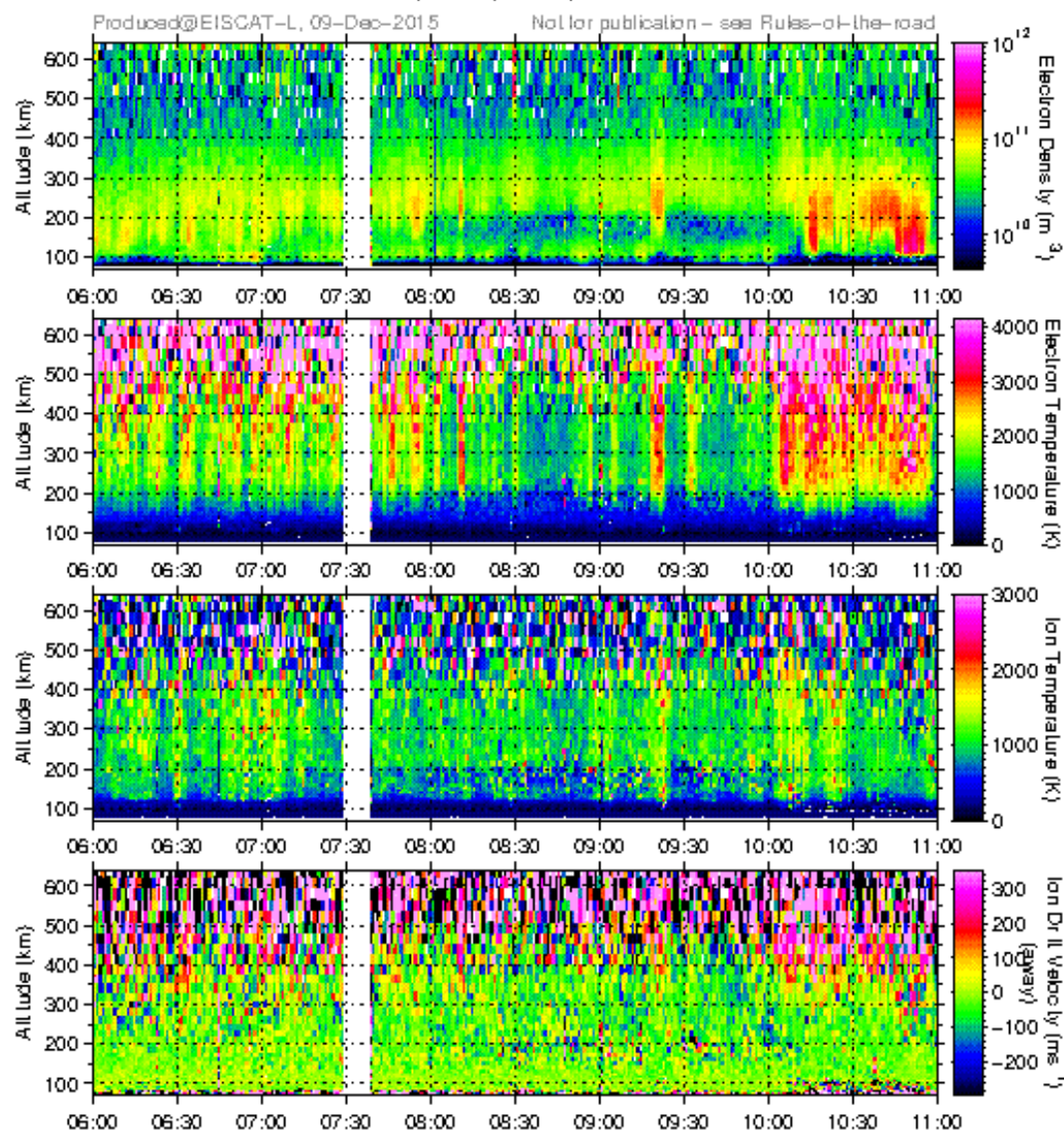




EISCAT Scientific Association

EISCAT SVALBARD RADAR

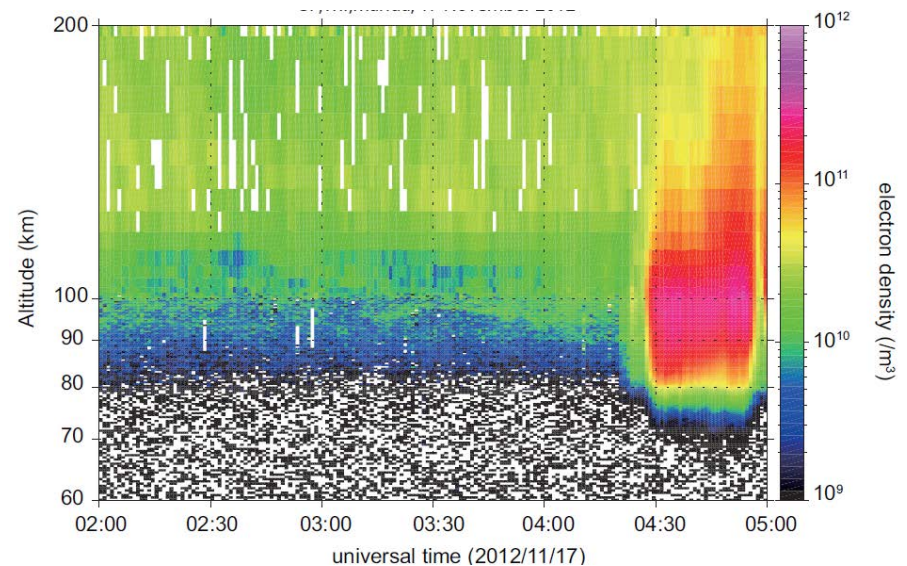
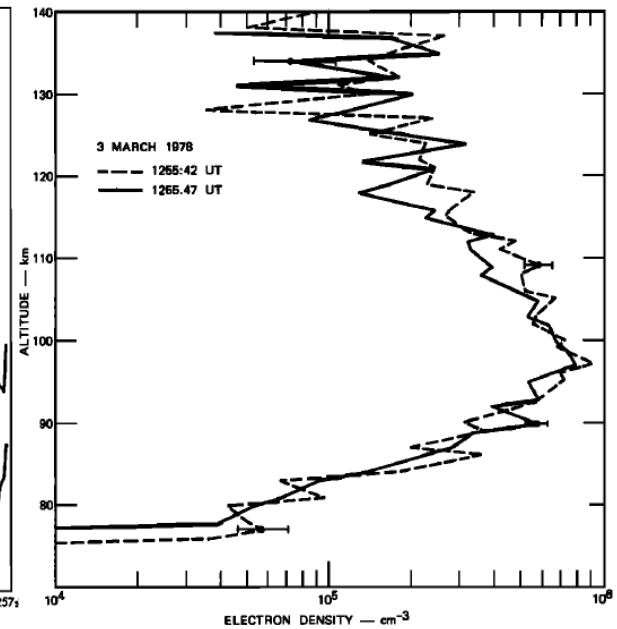
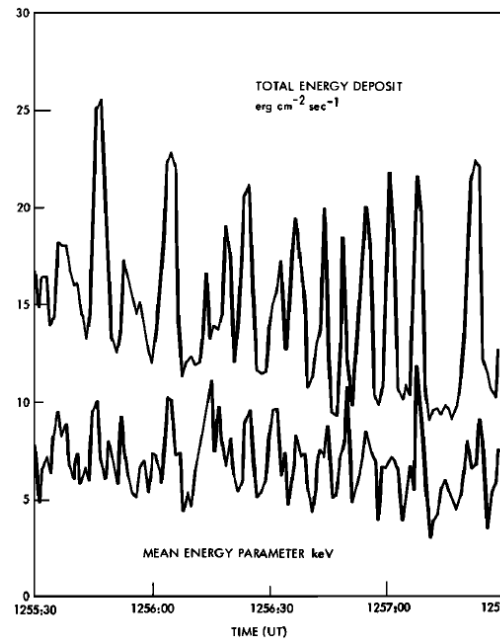
SP, 42ma, beata, 8 December 2015



RENU 2

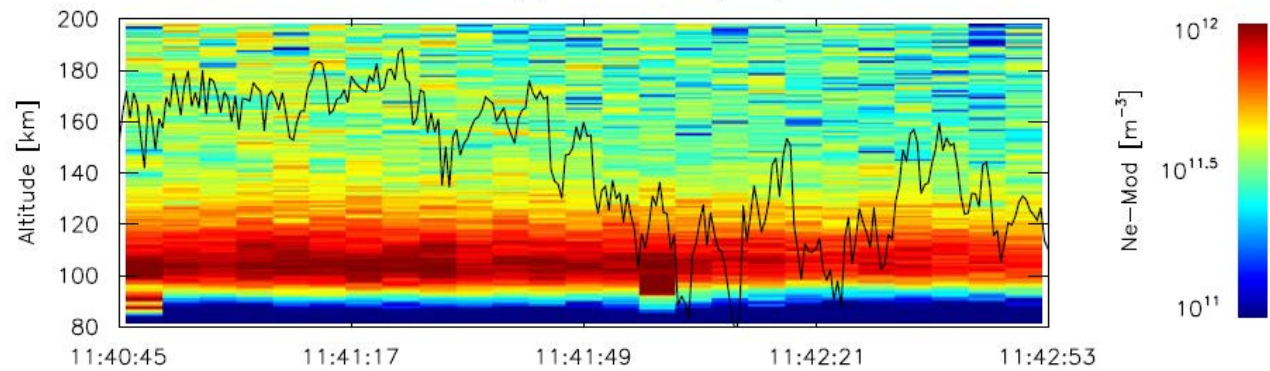
- Strong enhancement in electron temperature/density after 10:00
- Slight increase in ion temperature
- Visible vertical drift of ions associated with enhancement

- Chatanika, AK
 - Foster *et al* [1978]
 - Sears and Vondrak [1981] (figs. top right)
- EISCAT
 - Miyoshi *et al* [2015] (fig. bottom right)
 - Hosokawa and Ogawa [2015]
 - Turunen *et al* [2016]
- AMISR
 - Jones *et al* [2009]
 - Cosgrove *et al* [2010]

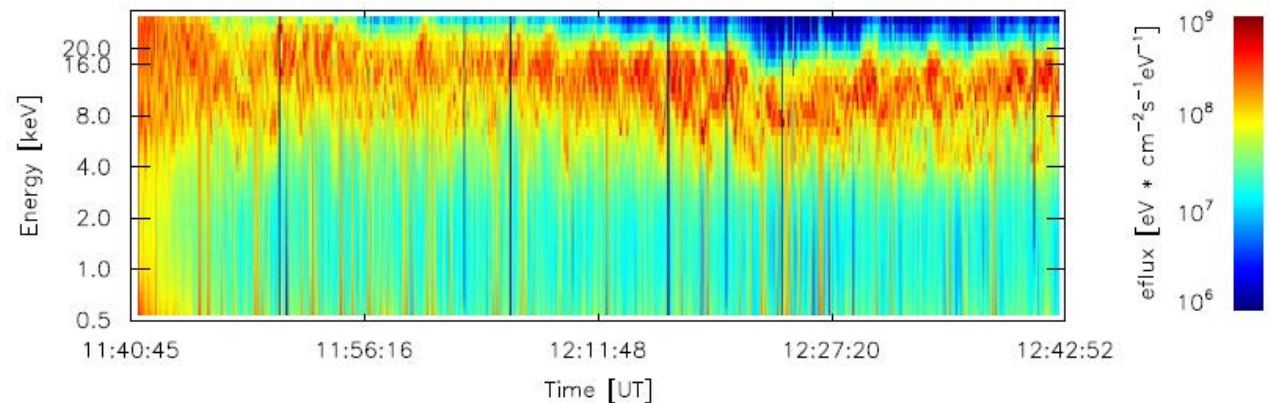


- First PFISR campaign
 - Spatial resolution: 1 km
 - Temporal resolution: 5 s
 - 480 μs long pulse interleaved with a 13 baud (10 μs) Barker code on two frequencies
- Invert electron density profile to determine energy distribution

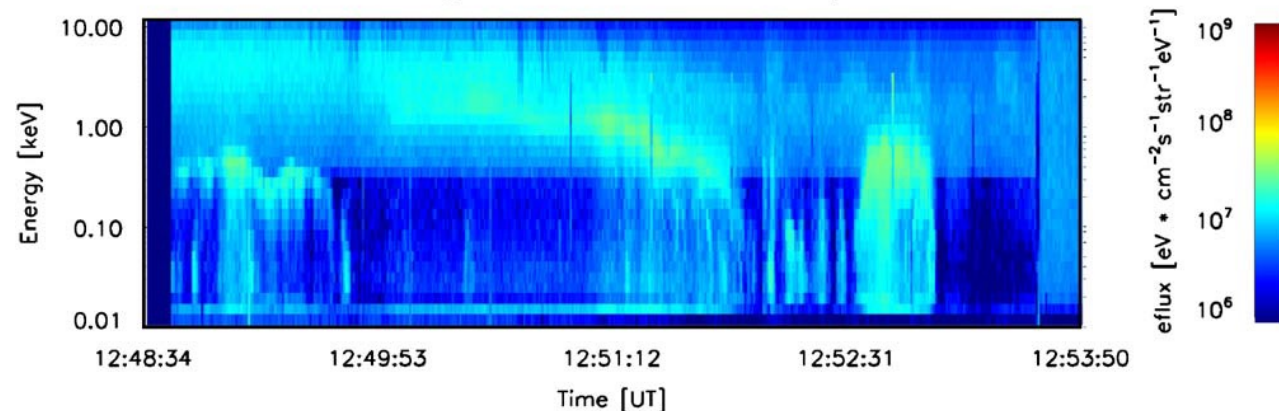
PFISR density profile: 2007/02/12



Energy distribution from PFISR density inversion



Differential Energy Flux from ROPA Main Payload HEEPS



Jones *et al* [2009]



Motivation



e-POP
FAI

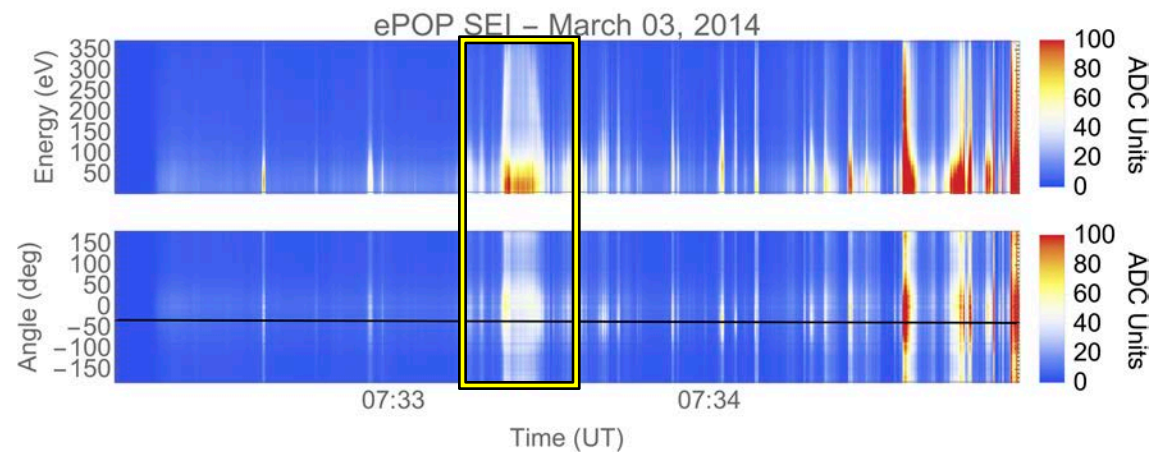
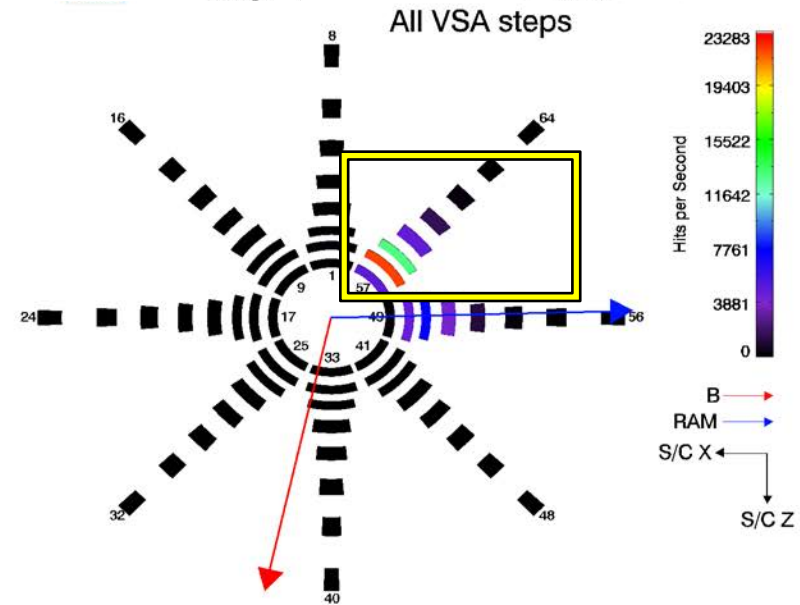


e-POP observation

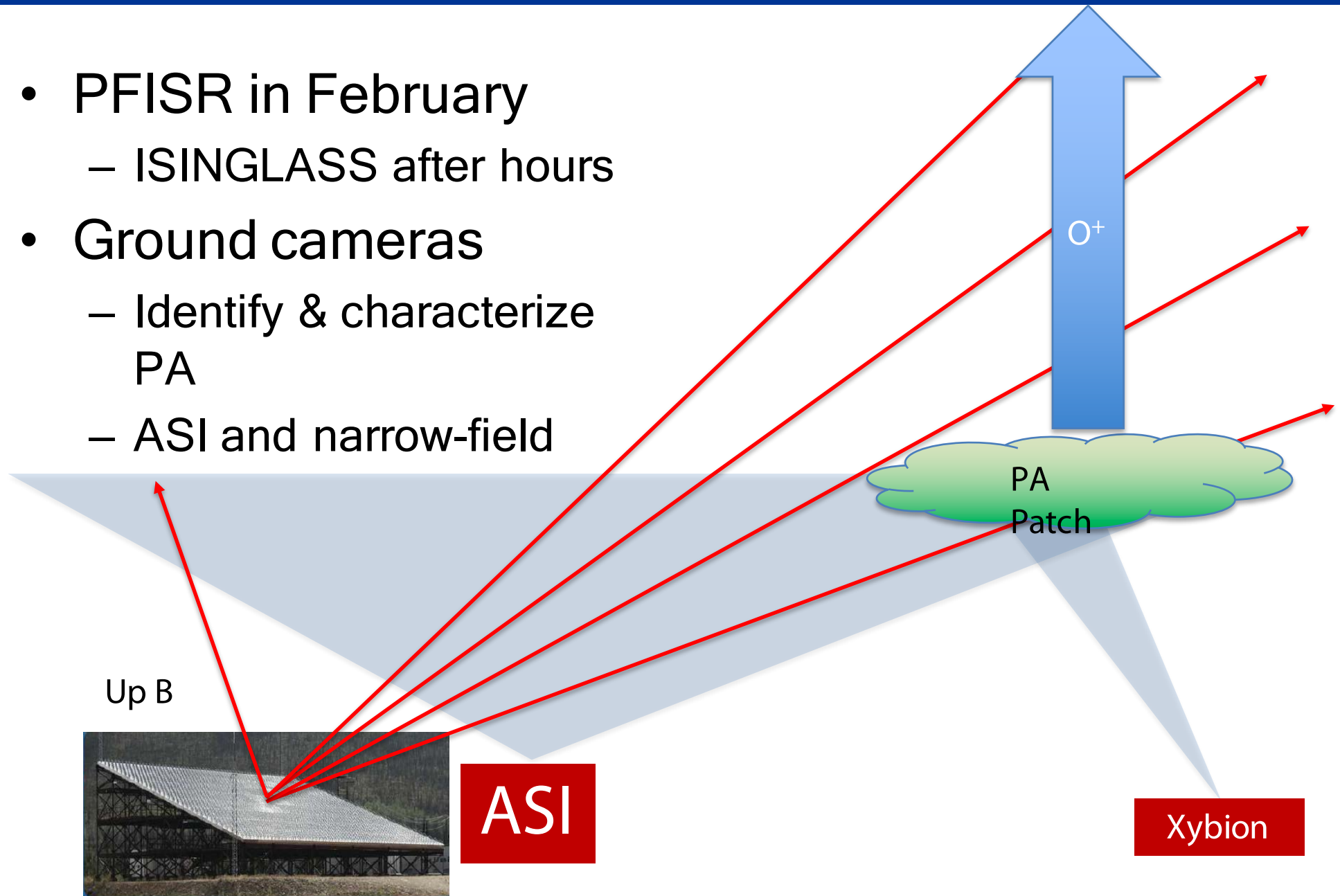
- IRM (ions)
 - Yellow box highlights upflowing O^+
 - Estimated velocity ~ 3 km/s
- SEI (electrons)
 - Low energy (~ 50 eV), isotropic electrons
 - thought to be backscatter



e-POP IRM March 03, 2014
07:33:30.17 - 07:33:31.11 UT



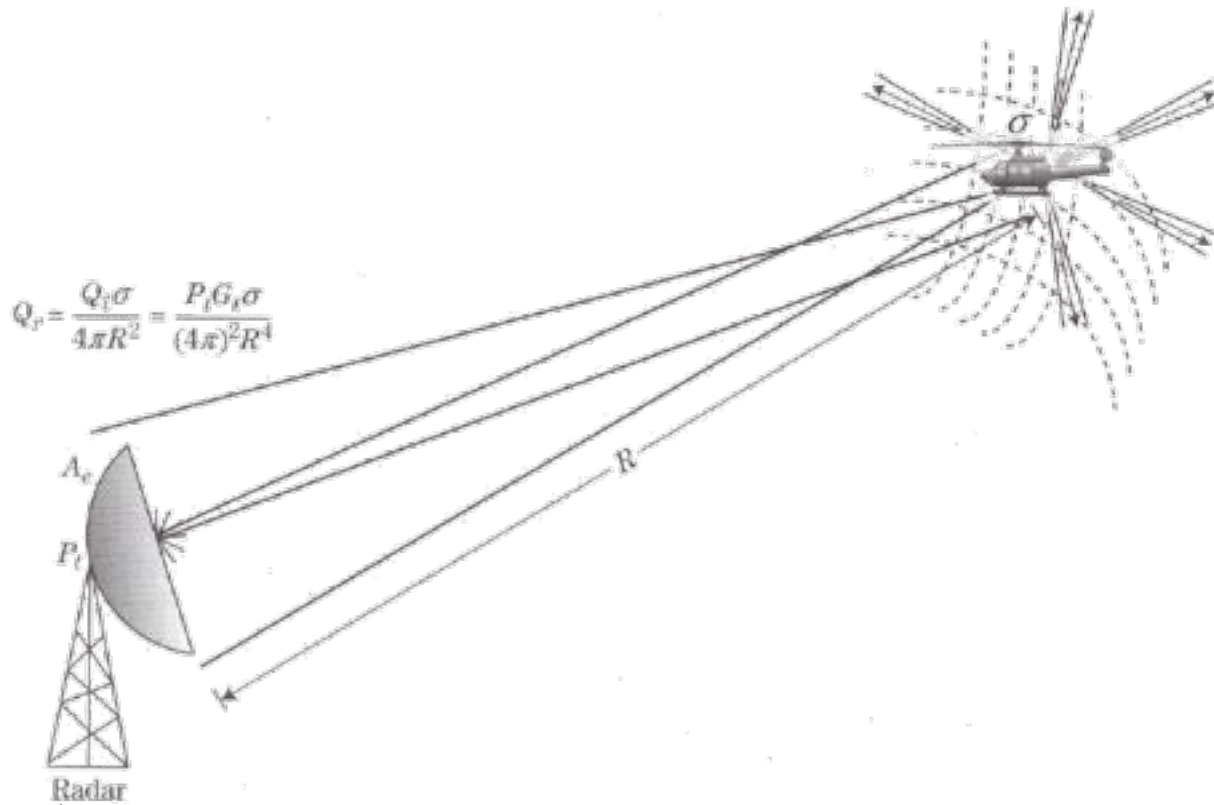
- PFISR in February
 - ISINGLASS after hours
- Ground cameras
 - Identify & characterize PA
 - ASI and narrow-field





Questions





$$Q_r = \frac{Q_t \sigma}{4\pi R^2} = \frac{P_t G_t \sigma}{(4\pi)^2 R^4}$$

