

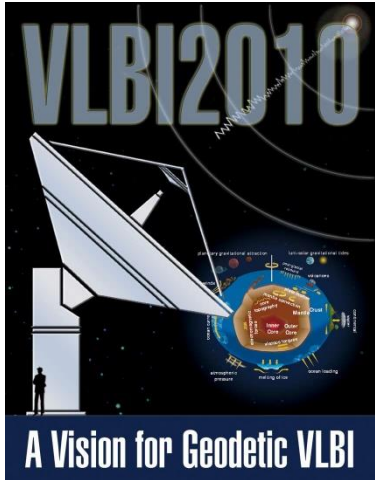


Realization Status of VGOS Infrastructure Buildout

Dirk Behrend

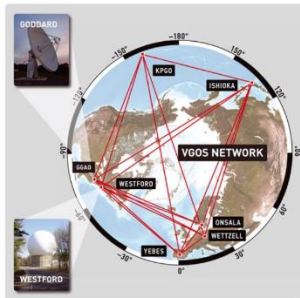
NVI, Inc./NASA GSFC, Greenbelt, MD, USA

11th IVS Technical Operations Workshop
May 3–5, 2021



- IVS Working Group 3: Final Report
 - https://ivscg.gsfc.nasa.gov/about/wg/wg3/IVS_WG3_report_050916.pdf
 - September 2005
 - Vision document

- VGOS Technical Committee (VTC): Progress Report
 - <https://ivscg.gsfc.nasa.gov/technology/vgos-docs/TM-2009-214180.pdf>
 - June 2009
 - Definition of specifications



- Radio Science: Technical Report
 - <https://doi.org/10.1029/2018RS006617>
 - September 2018
 - Overview of VGOS system



What is the IVS?



The **International VLBI Service for Geodesy and Astrometry (IVS)** is an international collaboration of organizations which operate or support Very Long Baseline Interferometry (VLBI) components:

- IVS inauguration was on **1 March 1999**.
- 83 permanent components supported by 41 institutions in 21 countries.
- ~300 Associate Members.

IVS is a recognized service of

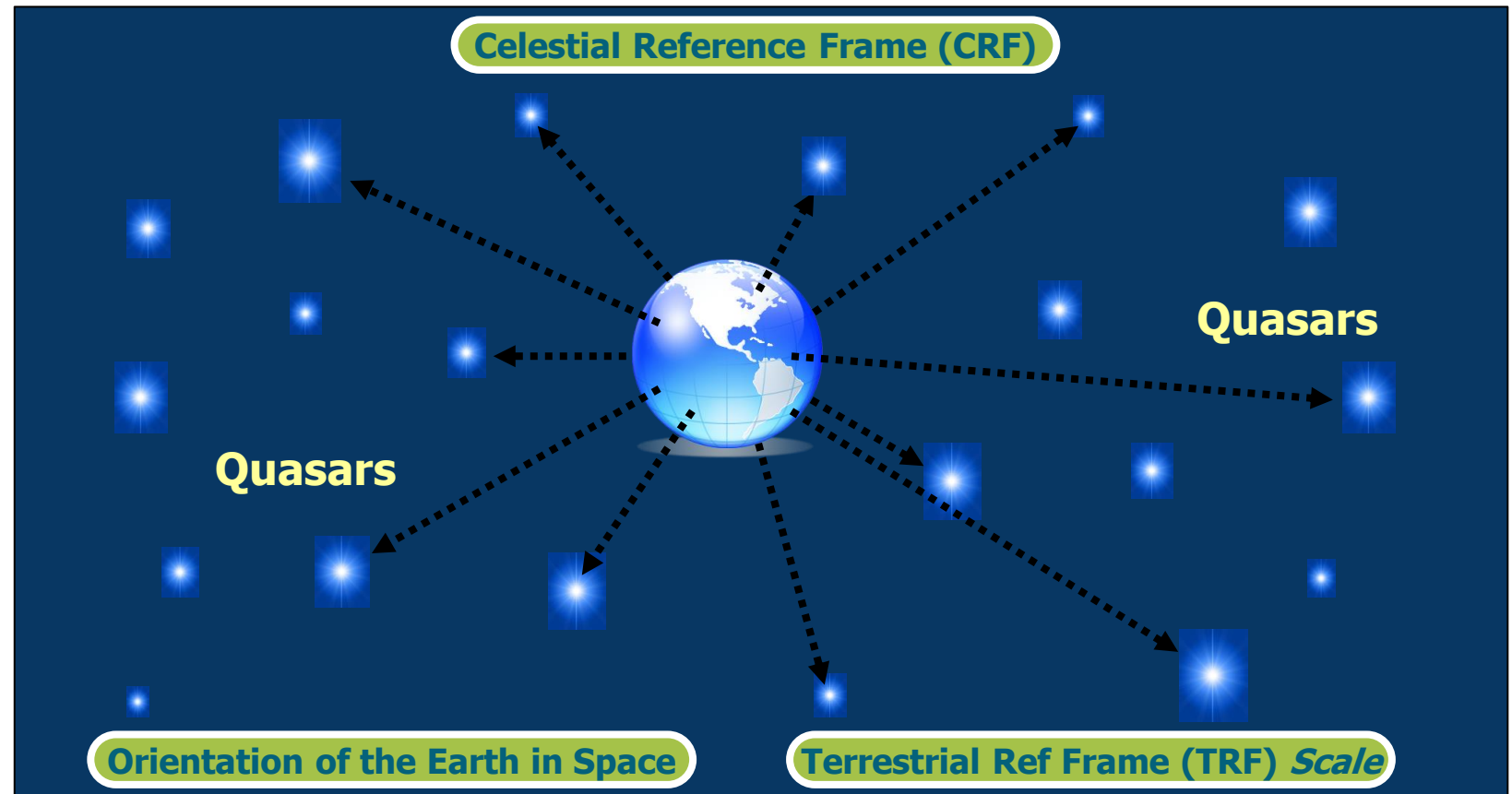
- **IAG** – International Association of Geodesy
- **IAU** – International Astronomical Union
- **WDS** – ISC World Data System

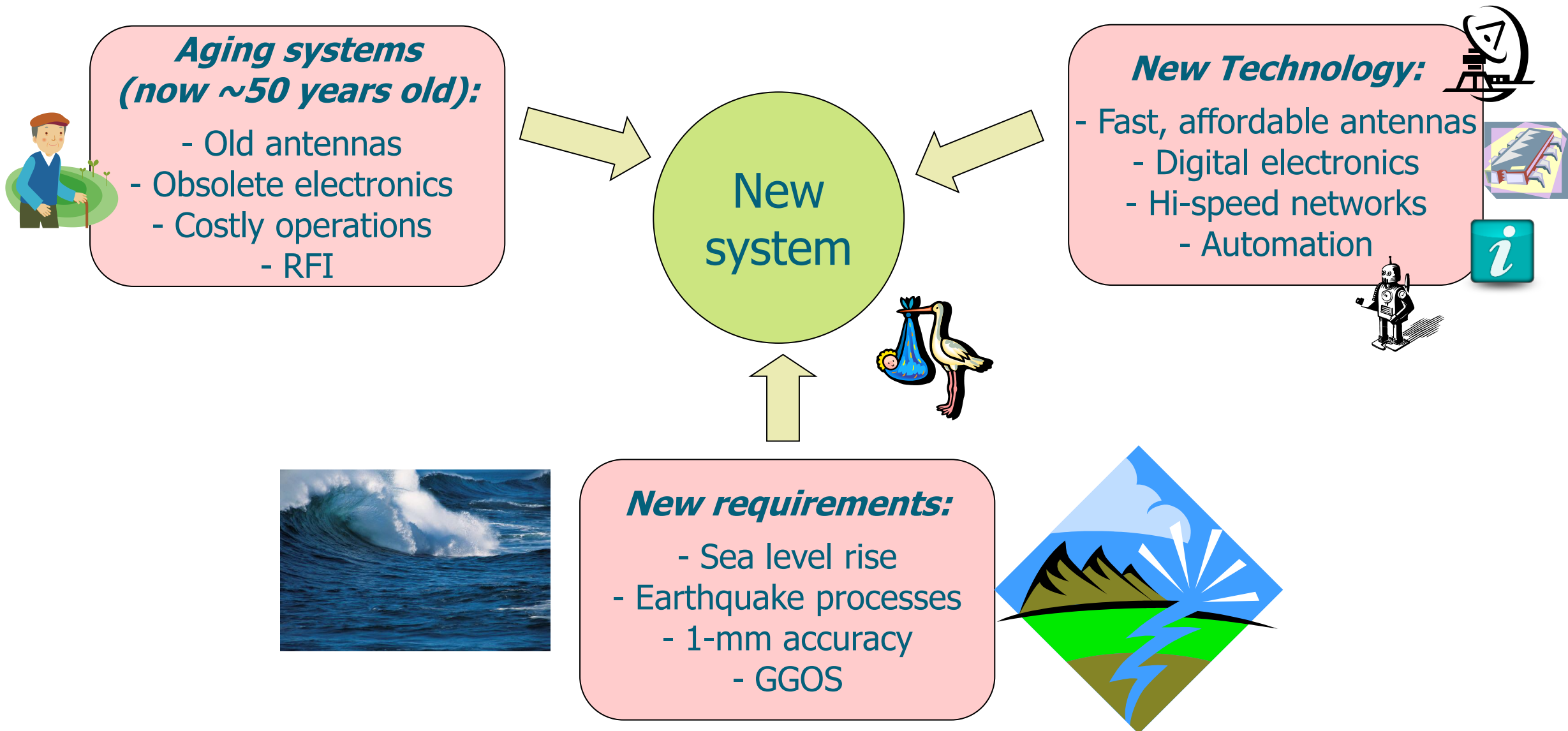
- Earth Orientation Parameters (EOP):
 - 24-hour sessions (all EOP)
 - 1-hour Intensives (UT1–UTC)

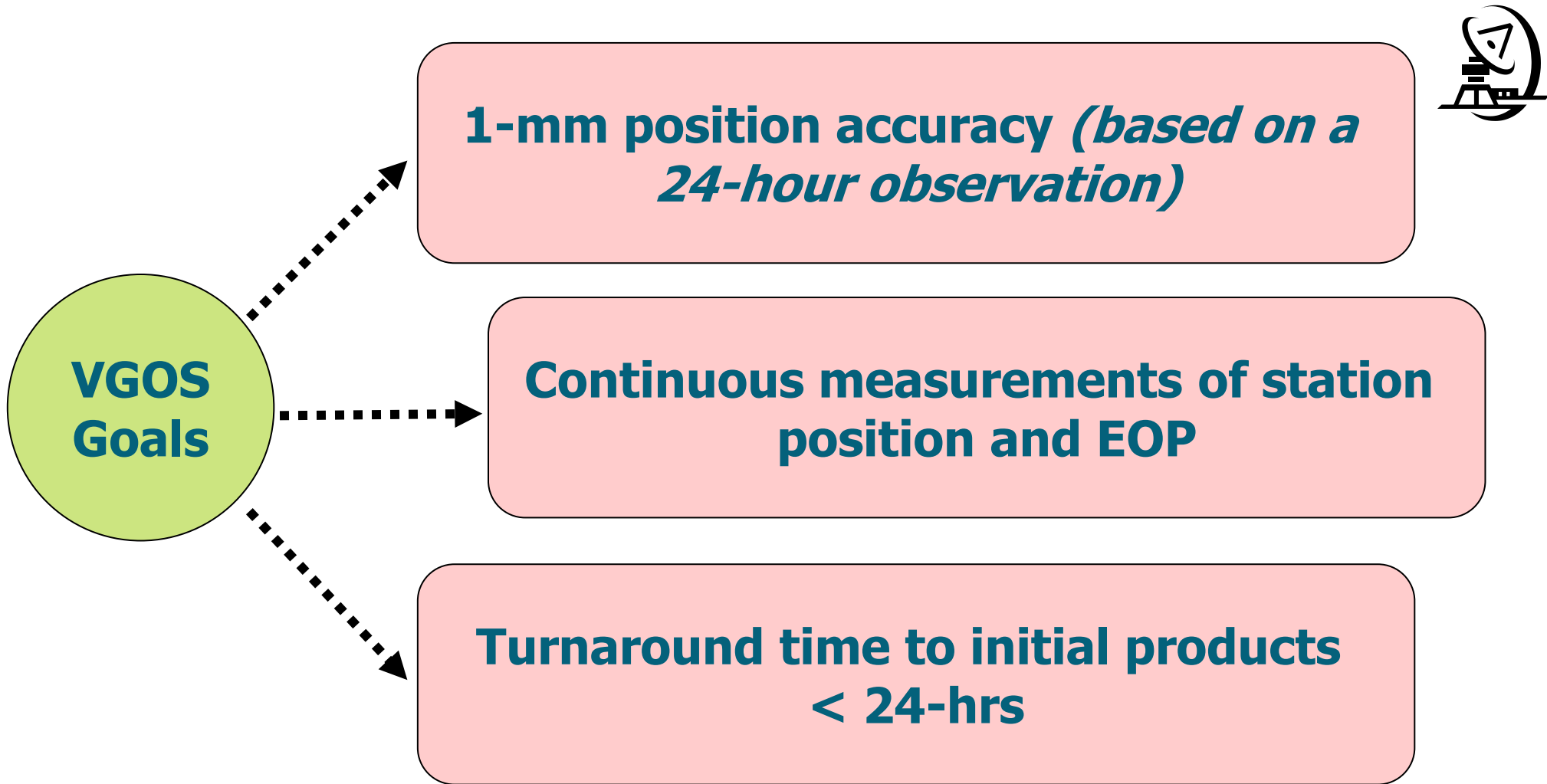
- Terrestrial Reference Frame (TRF)
- Celestial Reference Frame (CRF)



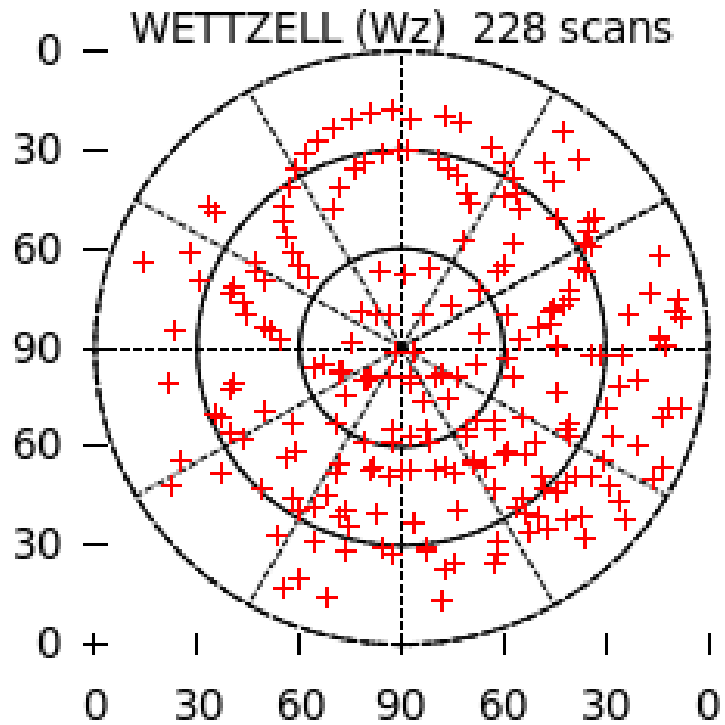
- Daily EOP + station coordinates (SINEX files)
- Tropospheric Parameters
- Baseline Lengths



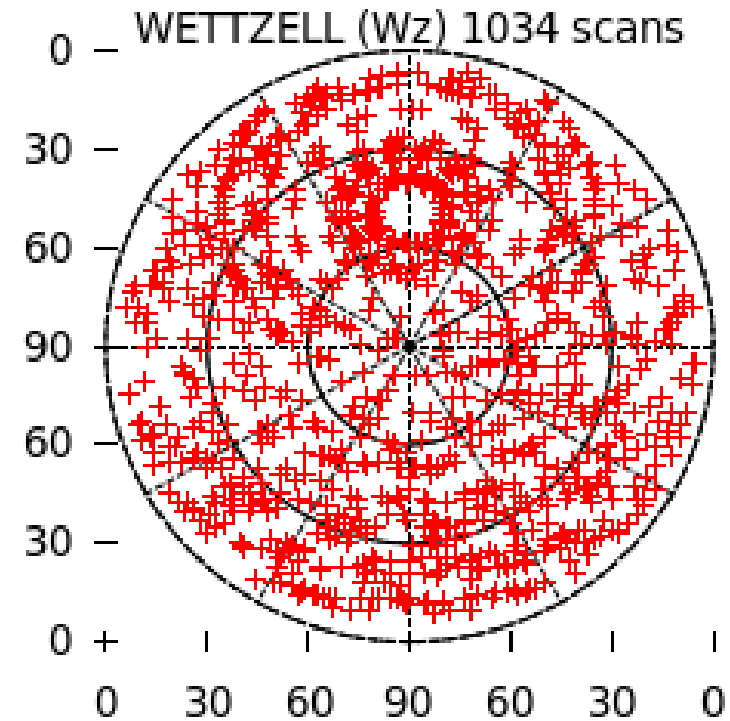




- dense sampling of local sky for optimal estimation of atmosphere parameters



Legacy S/X system



VGOS broadband system



VGOS vs. Legacy S/X: Comparison of Characteristics



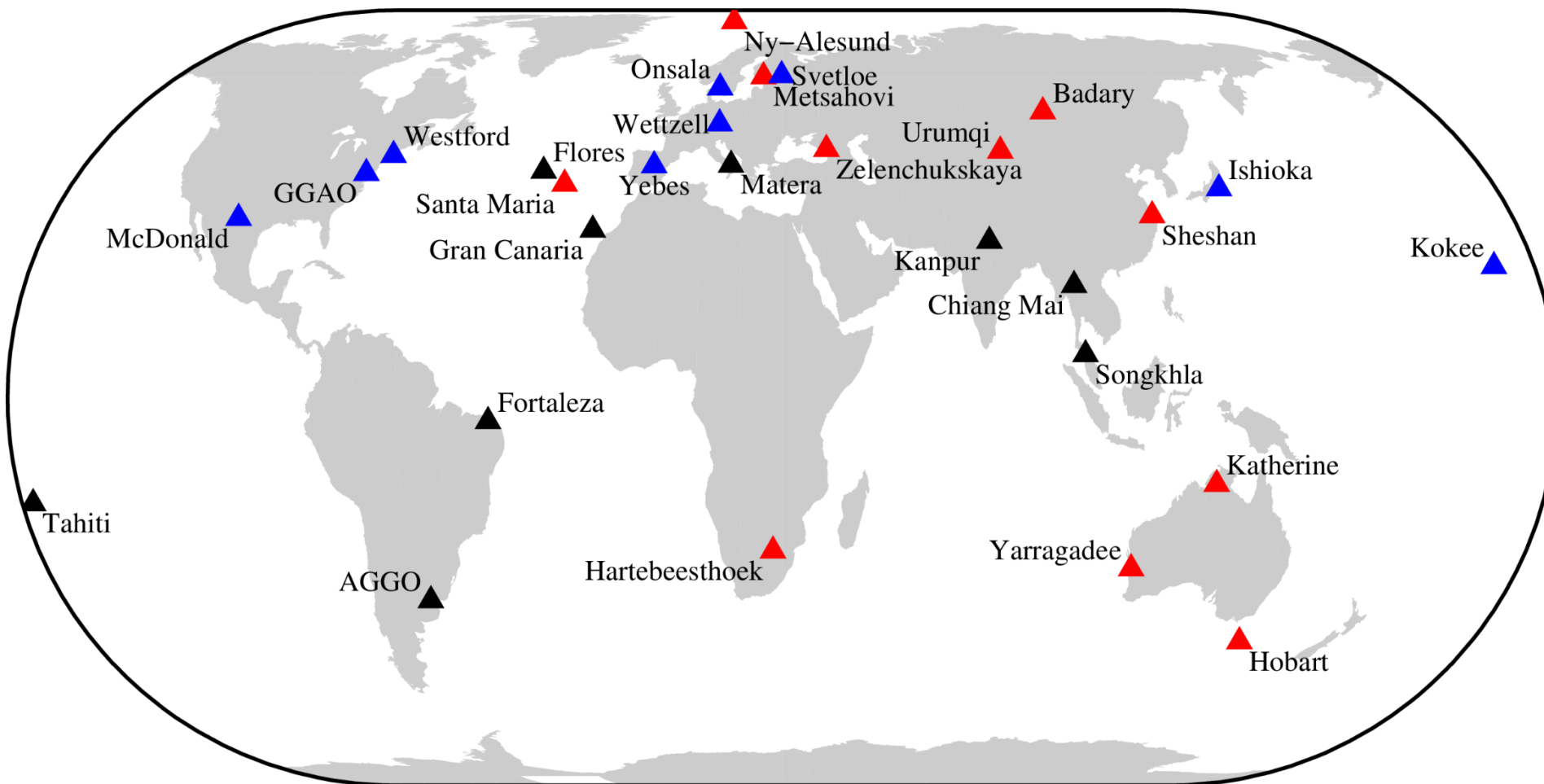
	Legacy S/X System	VGOS System	Benefit
Antenna size	5–100 m dish	12–13 m dish	reduced cost
Slew speed	~20–200 deg/min	≥ 360 deg/min	more observations for troposphere
Sensitivity	200–15,000 SEFD	≤ 2,500 SEFD	more homogeneous
Frequency range	S/X band [2 bands]	~2–14 GHz [1 broadband w/ 4 bands]	increased sensitivity, data precision
Recording rate	128, 256, 512 Mbps	8, 16, 32 Gbps	increased sensitivity
Data transfer	usually e-transfer, some ship disks	e-transfer, ship disks when required	
Signal processing	analog/digital	digital	stable instrumentation

Onsala Skyline



Courtesy R. Haas

Evolution of the VGOS Network



▲ operational
 ▲ antenna built, signal chain work
▲ in planning stage



VGOS Observing Sessions



VGOS observing in 2019:

- VGOS Test (VGOS-T) sessions
 - One 24-hour session every 2 weeks, correlation at MIT Haystack Observatory
 - Databases made available at IVS data centers (e.g., CDDIS)

VGOS observing in 2020:

- VGOS Operational (VGOS-O) sessions
 - One 24-hour session every 2 weeks, correlation at MIT Haystack Observatory
 - Officially operational sessions, databases at IVS data centers
- VGOS Intensive (V2) sessions
 - (Starting in late February) one 1-hour session every 2 weeks, alternating with VGOS-O sessions
 - Correlation at MIT Haystack Observatory, databases at IVS data centers



VGOS Correlation Verification



Developer:

- MIT Haystack Observatory
- Correlation and fringe fitting process based on DiFX software correlator

Knowledge transfer:

- Correlation workshop in May 2019
 - Transfer procedures, software versions, lessons learned to other centers to increase VGOS observing frequency
- Follow-up workshop in 2020 (postponed due to COVID-19)
- Correlation comparisons using benchmark data sets (1-hour, 24-hour)

Verification of correlators:

- Bonn, Shanghai, USNO, and Vienna (plus Tsukuba and Onsala)
 - Each correlator has unique data transport challenges
 - Accepts disk modules, e-transfer only, limited network capacity
- Hands-on “blind-test” correlation with 1-hour VGOS Intensive session
 - Starting at raw data level
 - Verified that results agree within margin of errors
 - Needed iteration



VGOS Correlation Verification

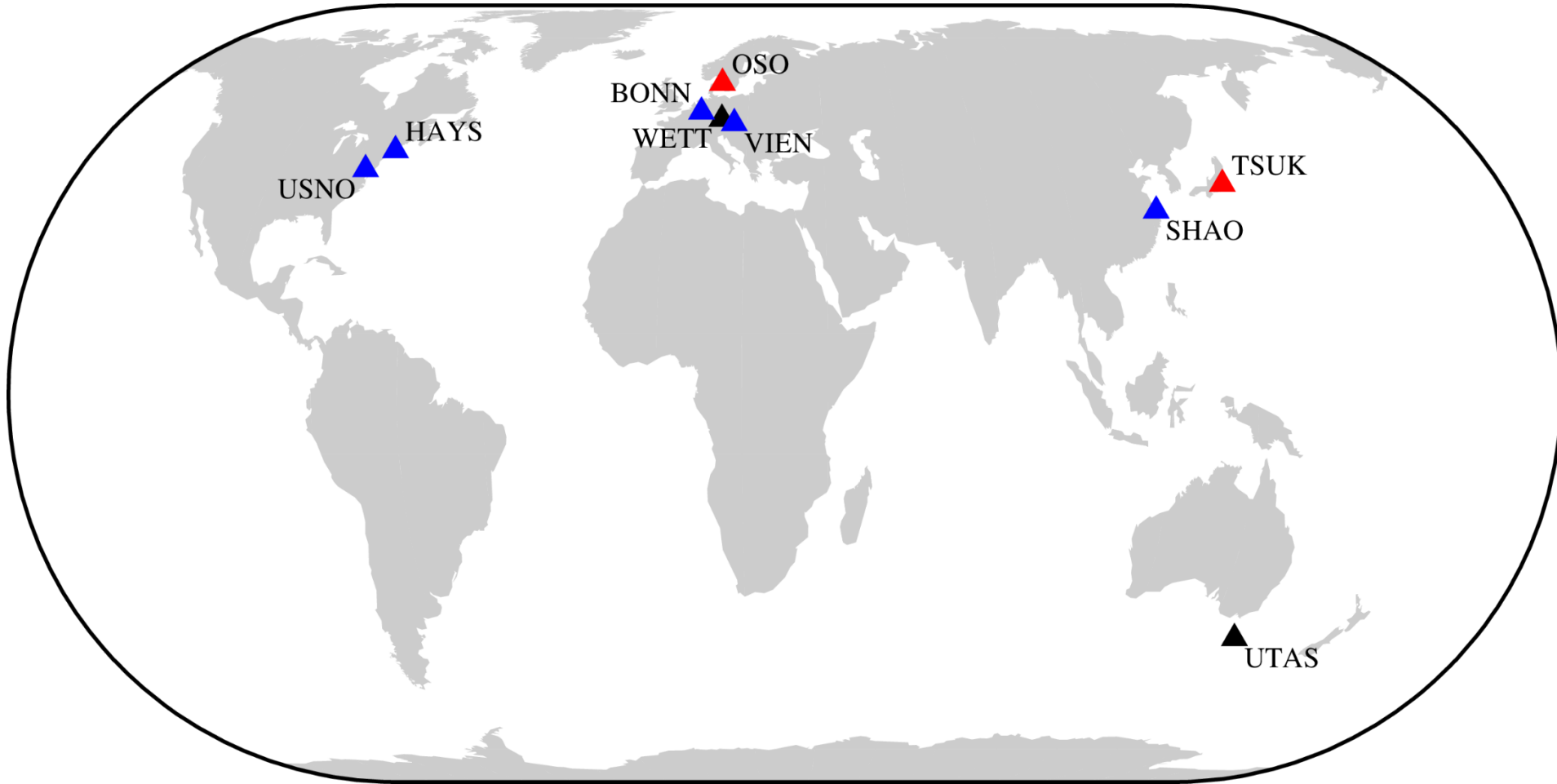


- Final verification w/ operational 24-hour VGOS session (VGOS-O: VO0009)
 - Bonn end-to-end from raw level
 - All other correlators start at post-correlation level (due to challenges w/ data transport of huge raw data)
 - Verified and validated VGOS correlation end-to-end process
 - Bonn, USNO started w/ operational sessions in 2020; VIEN, SHAO followed in 2021
- All verification results are published in [Haystack memo series](#)

“Mixed-mode” correlation:

- Simultaneous observing with VGOS and S/X stations in same network
- Important for tying the VGOS frame to the S/X frame
- Three R&D sessions observed in summer of 2020; two now correlated and contributed to ITRF2020
- Correlation VGOS-VGOS, S/X-S/X, and VGOS-S/X
- Mixed-mode correlator workshop directly following TOW2021

Rollout of VGOS Correlation Capabilities



▲ operational ▲ under verification
▲ future correlation center



Data Storage and Transport



Data storage requirements:

- Legacy S/X VLBI: ~2000 TB/year
- Typical 24-hour VGOS session (2020):
 - 8 stations
 - 50 TB/day/station of raw data
 - ~400 TB/day
- Network size to grow to 16–20 stations
 - for 20 stations: ~1000 TB/day
 - full year (24/7/365): ~360 PB/year

Data transport (electronic transfer):

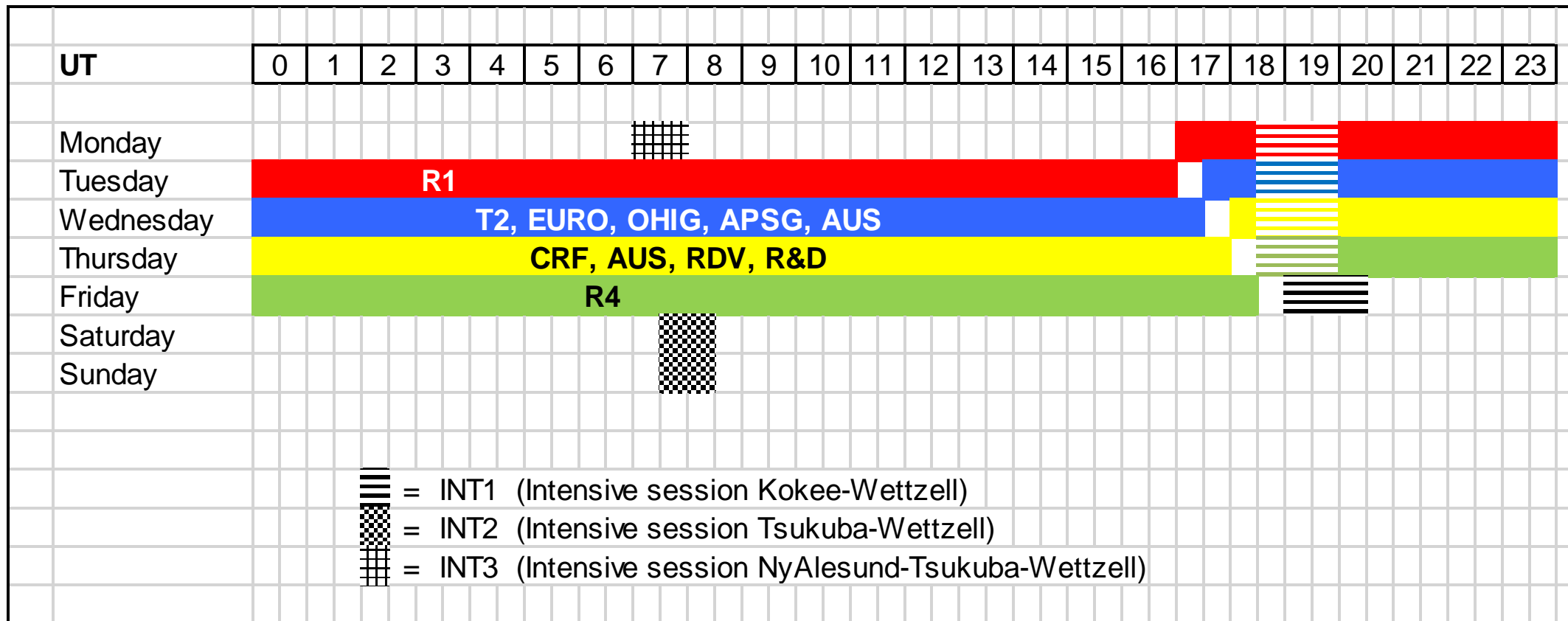
- Required network data rates:
 - each site: 5.6 Gbps
 - correlator: 134 Gbps

Station	Bandwidth now (sustained)	Transfer time for 50 TB of data
GGAO	1 Gbps	4.75 days
Westford	1 Gbps	4.75 days
Wettzell	4 Gbps	1.2 days
Yebeş	5 Gbps	0.95 days
Ishioka	1 Gbps	4.75 days
Kokee Park	0.1 Gbps	47.50 days
Onsala	6 Gbps	0.8 days
McDonald	0.3 Gbps	15.8 days

Note:

Kokee was recently upgraded to 1 Gbps

➤ Typical weekly layout for IVS observing sessions



➤ about 180 sessions per year, 3.5 sessions per week

- Expected weekly observing coverage for VGOS (after 2020)

UT	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23
Monday																								
Tuesday																								
Wednesday																								
Thursday																								
Friday																								
Saturday																								
Sunday																								
	Constant observation with 16+ station network Individual stations have maintenance days																							

- 365 sessions per year, 7 sessions per week (24/7/365)



Analysis: VGOS Results for UT1–UTC



Comparison of VGOS Intensive and legacy S/X Intensive results w.r.t. IERS Bulletin B:

- The VGOS-B data show both the smallest uncertainties, a small bias, as well as the best agreement in terms of RMS and STD.
- Based on 11 sessions only, but very encouraging



Series	σ_{mean}	σ_{median}	RMS	Bias	STD
All BKG INT1/INT2	14.8	14.2	32.9	14.3 ± 3.4	29.6
All GSF INT1/INT2	15.1	13.2	28.3	-2.0 ± 3.2	28.2
All USN INT1/INT2	14.6	12.9	28.4	8.5 ± 3.1	27.1
All GIS INT2	13.5	9.2	33.6	10.2 ± 6.7	32.0
All IAA INT1/INT2	15.0	12.2	27.8	5.0 ± 3.6	27.3
All OSO INT1/INT2	16.5	15.4	31.0	4.5 ± 3.6	30.6
OSO VGOS-B	4.5	4.2	23.2	-3.8 ± 7.2	22.9
Simultaneous OSO INT1	16.0	14.2	28.4	-0.8 ± 9.0	28.3
Simultaneous BKG INT1	17.3	16.1	32.9	8.4 ± 10.1	31.8
Simultaneous GSF INT1	18.8	16.8	24.2	-7.6 ± 7.3	23.0
Simultaneous USN INT1	14.3	14.4	27.9	5.4 ± 9.1	27.4

From: Haas et al. Earth, Planets and Space (2021) 73:78

- VGOS network expansion to 24+ stations in next few years
- Use numerous correlators to go to higher cadence VGOS sessions
- Increase data storage and data transfer capacities at stations and correlation facilities
- Tie VGOS TRF to legacy S/X TRF using mixed-mode sessions (VGOS–S/X)
- Further avenues of improvement:
 - instrumentation development (e.g., bandwidth doubling)
 - atmosphere modeling
 - radio source structure
- Transition IVS production system from legacy VLBI to VGOS
- Eventual 24/7 observing

Ny Ålesund Twin Telescope



Thank you!

*Inauguration event
on 6 June 2018*